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# On the Measurement of Geostrophic Ocean Currents by (NADIR) Satellite Altimeter

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With the assumptions of accurately known marine geoid, satellite orbit, ionospheric and atmo-					
spheric propagation corrections, altimeter range precision, and dynamic components of sea-level					
deviations due to tides, barometric pressure, wind set-up, and storm surges, orbit inclination of 108° is shown to be the best choice among the three selected orbit inclinations (93°, 98°, and 108°) for the					
detection of geostrophic current vectors above 20° latitude. The second choice is 98° orbit inclination.					
The relationship shown as	de. The second choice is 38 orbit inclination,				
V <sub>R</sub> = C	v				
<b>∀R</b> - <b>C</b>	(Continues)				

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### ON THE MEASUREMENT OF GEOSTROPHIC OCEAN CURRENTS BY (NADIR) SATELLITE ALTIMETRY

#### 1.0 INTRODUCTION

The following discussion considers the problem of measuring geostrophic ocean current velocities with a single-beam, pulse-limited, nadir-looking satellite radar altimeter similar to those on SKYLAB, GEOS, and SEASAT. It is assumed that sufficient information is available to accurately determine the instantaneous, local, dynamic ocean heights as measured with respect to the marine geoid. These measurements require precise knowledge of the local geoid, satellite orbit, and non-current-related factors affecting measurement of ocean surface topography. It is also further assumed that the differences in satellite orbit slope errors at orbit crossings are negligible in their effects on the derivation of the local dynamic ocean surface slopes.

Principal altimeter measurement corrections include uncertainties of the order of 20 cm for both atmospheric propagation and ionospheric effects [Saastamoinen, 1971; Davies, 1965; Thomason, et al, 1979]. Dynamic processes causing local deviations of the ocean surface topography from the marine geoid include: tides (30 cm open-ocean, 15 m coastal) [Neumann and Pierson, 1966; Doodson, 1958; Stoker, 1957]; barometric pressure gradients (up to 1 m); wind set-up and storm surges (up to 10 m in coastal regions) [Neumann and Pierson, 1966]; and density gradients associated with current systems (1 m associated with strong boundary currents such as the Gulf Stream and the Kuroshio Current) [Stommel, 1966; Stommel and Yoshida, 1972]. The measurement problem is illustrated in Figure 1.1.

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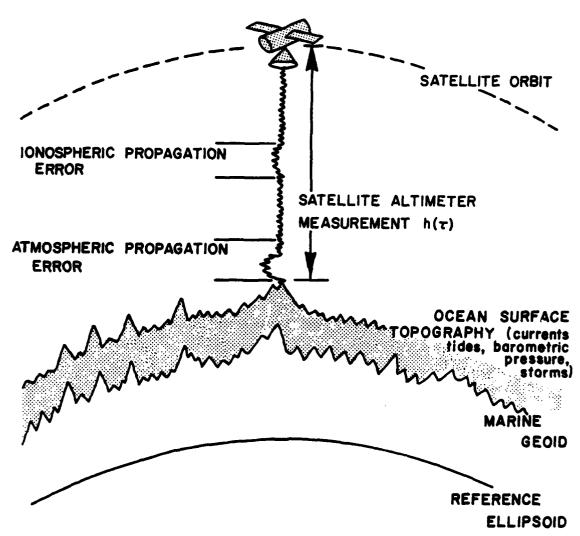


Fig. 1.1 - Configuration of altimeter measurement

### 2.0 MEASUREMENTS/ASSUMPTIONS

- 2.1 Assumptions. For the following discussions it is assumed that:
  - (1) The marine geoid is accurately known.
  - (2) The satellite orbit is accurately determined.
  - (3) Data are available for accurate determination of ionospheric and atmospheric propagation corrections to satellite altimeter range measurements.
  - (4) Satellite radar altimeter range measurement precision has been determined.
  - (5) Dynamic components of sea-level deviations due to tides, barometric pressure, wind set-up, and storm surges are accurately determined from models and other data input.
- 2.2 Measurements. For the following discussions, it is presumed that the assumptions above are satisfied and that the measurement to be considered is the current-related dynamic ocean surface topography within the precision limit of the radar altimeter measurement. Using time (t) as the parameter to define position along the sub-satellite track, the current-related dynamic ocean surface topography can be expressed as:

$$D(t) = O(t) - [h(t) - \varepsilon_{I}(t) - \varepsilon_{A}(t)] - [G(t) + T(t) + B(t) + W(t)]$$
where:

- O(t) is the satellite orbit,
  - h(t) is the radar range measurement,
  - $\varepsilon_{\tau}(t)$  is the ionospheric propagation correction,
  - $\varepsilon_{\Lambda}(t)$  is the atmospheric propagation correction,
  - G(t) is the marine geoid,
  - T(t) is the tide,
  - B(t) is the topographic variation due to barometric pressure, and
  - W(t) is the topographic variation due to wind set-up and storm surges.

Both O(t) and G(t) are defined with respect to the center of mass of Earth. The determinations of: O(t),  $\epsilon_{\bar{I}}(t)$ ,  $\epsilon_{\bar{A}}(t)$ , G(t), T(t), B(t), and W(t) are assumed to be absolutely correct.

#### 3.0 DETERMINATION OF GEOSTROPHIC OCEAN CURRENT

Following Sverdrup, et al. [1946] ocean currents may be divided into three groups: (1) currents that are related to the distribution of density in the sea (barotropic and baroclinic geostrophic currents), (2) currents that are caused directly by the stress that the wind exerts on the sea surface (wind driven currents), and (3) tidal currents and currents associated with internal waves. In accordance with the above-stated assumptions, the following discussion concerns only the geostrophic (density-gradient) currents.

The geostrophic current is determined by a balance between the Coriolis force (acceleration due to the earth's rotation and the pressure gradients (slope of the isobaric surfaces) in the ocean. The numerical value of the ocean-current velocity for a given slope of an isobaric surface is given by:

$$V_{R}(\phi) = g \frac{\tan S_{R}}{2\omega \sin \phi}$$
 (3.1)

where:

 $V_{\mathbf{p}}(\phi)$  = geostrophic current speed in m/sec at Latitude  $\phi$ ,

g = acceleration of gravity,

tanSp = slope of the isobaric surface,

S<sub>R</sub> = angle of the dynamic ocean surface to the local horizontal plane,

 $\omega$  = angular velocity of rotation of earth, and

\$\phi\$ = geographic latitude which is positive and negative
for northern and southern hemispheres, respectively.

In the northern hemisphere ( $\phi$ >0), the current direction is 90° to the left of the direction of positive slope. In the sourthern hemisphere ( $\phi$ <0), the current direction is to the right.

Taking:  $g = 9.80 \text{ m/sec}^2$ , and  $\omega = 7.27 \text{x} 10^{-5} \text{ rad/sec}$  in Equation (3.1)

$$\mathbf{V}_{\mathbf{R}}(\phi) = 6.74 \times 10^4 \frac{\tan S_{\mathbf{R}}}{\sin \phi} \tag{3.2}$$

At 45° latitude, for example:

$$V_R(45^\circ) = 9.53 \times 10^4 tanS_R$$
 (3.2a)

and if  $tanS_R = 10^{-5}$  (10 cm height difference over 10 km):

$$V_p(45^\circ) \simeq 1 \text{ m/sec} \simeq 2 \text{ kts.}$$
 (3.2b)

If the surface slope is expressed in dimensions of sec, the expression for geostrophic current speed becomes, with  $\omega = 15$  sec/sec,

$$V_{p}(\phi) = 0.3268 \tan S_{p}/\sin \phi \tag{3.3}$$

at 45° latitude, for example,

$$V_p(45^\circ) = 0.462 \, \tan S_p$$
 (3.3a)

and if tanS<sub>R</sub>= 1 sec

$$V_p(45^\circ) \approx 0.5 \text{ m/sec} \approx 1 \text{ kt.}$$
 (3.3b)

## 4.0 GEOSTROPHIC CURRENT INFERRED FROM ALTIMETER MEASUREMENTS AT THE CROSSING POINT

It is well recognized that a single-beam nadir radar altimeter cannot provide measurement of the geostrophic current vector from a single along-track surface slope. The current derived from a single along-track surface slope is, at best, the across-track current component, let alone the true geostrophic current. The scheme discussed in Section 6.0 gives reasonable recourse for short term descriptions of the geostrophic current system.

Nevertheless, the true local geostrophic current velocities can be inferred from altimeter measurements at the satellite subtrack crossing points with the additional assumption: the geostrophic

current system has not experienced significant changes during the time interval which is required for the satellite to complete these crossings. In the case of SEASAT-A, it takes the satellite less than 11 hours to complete a crossing pair [Cutting, et al, 1977]. Since most of the time scales for geostrophic phenomena in the ocean are longer than 11 hours, the assumption is valid, most of the time. From the two local dynamic ocean surface slopes derived from the altimeter measurements made at the crossings the local dynamic ocean surface plane can be constructed from which its true slope vector can be evaluated for this local plane [Sokolnikoff,1951]. Then the geostrophic current velocities may be determined from Equation (3.1) with the known directions.

Let  $\overline{\text{OA}}_1$  and  $\overline{\text{OA}}_2$  be these two slope vectors without separation at origin O in a Right Hand Rectangular Cartesian Coordinates as shown in Figure 4.1 where

$$A_1 = (t_1 \cos(\pi - \psi/2), t_1 \sin(\pi - \psi/2), t_1 s_1).$$

$$A_2 = (t_2 \cos(\pi + \psi/2), t_2 \sin(\pi + \psi/2), t_2 s_2),$$

and  $t_i$  and  $s_i$  are the along-track horizontal distance and measured slope, respectively. The intersection angle  $\psi$  between the ascending and the descending orbits is bisected by the x-axis which is, of course, in the local east-west direction. These intersection angles with respect to latitude for selected orbit inclinations are shown in Figure 4.2.

The plane which contains points  $A_1$ ,  $A_2$ , and the origin 0 has the general form of

$$a_0x + b_0y + c_0z + d_0 = 0$$
 (4.1)

with the origin 0 in the plane  $d_0 = 0$  or

$$z = ax + by (4.2)$$

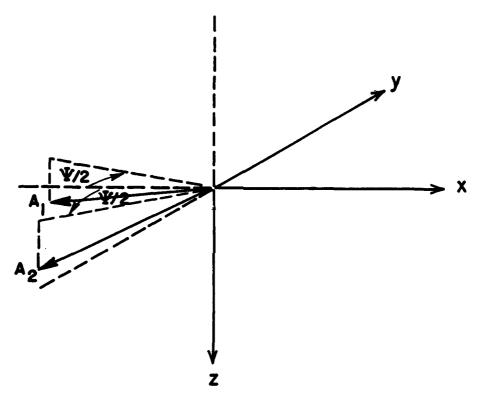


Fig. 4.1 - Rectangular Cartesian coordinates with Jacobian = 1

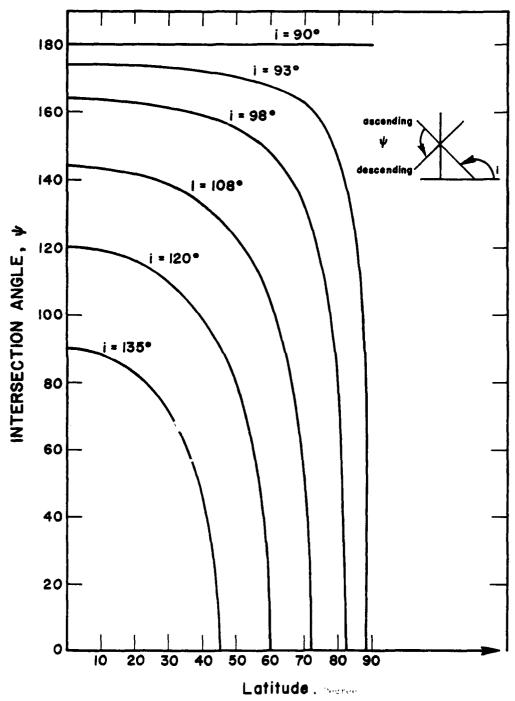


Fig. 4.2 - Intersection angle with respect to latitude for orbit inclinations of  $90^{\circ}$ ,  $93^{\circ}$ ,  $98^{\circ}$ ,  $108^{\circ}$ ,  $120^{\circ}$  and  $135^{\circ}$ 

Substituting the coordinates of points  $A_1$  and  $A_2$  into Equation (4.2)

$$\cos(\psi/2) \ a - \sin(\psi/2) \ b = - s_1$$
 (4.3)

$$\cos(\psi/2) \ a + \sin(\psi/2) \ b = -s_2$$
 (4.4)

Therefore, from Equations (4.3) and (4.4)

$$a = -\frac{s_1 + s_2}{2\sin(\psi/2)} \tag{4.5}$$

$$b = \frac{s_1 - s_2}{2\sin(\psi/2)} \tag{4.6}$$

With a and b known, Equations (4.2), the plane which contains points  ${\bf A}_1$ ,  ${\bf A}_2$ , and the origin 0 is

$$\frac{s_1 + s_2}{2\cos(\psi/2)} \times -\frac{s_1 - s_2}{2\sin(\psi/2)} \quad y + z = 0$$
 (4.7)

The direction cosines of the normal to this plane are

$$\frac{s_1 + s_2}{2\cos(\psi/2)}$$
 ,  $-\frac{s_1 - s_2}{2\sin(\psi/2)}$  , 1 (4.8)

and the angle,  $\gamma_n$ , of the normal to the z-axis is defined as

$$\cos \gamma_{n} = \left\{ \left[ \frac{s_{1} + s_{2}}{2\cos(\psi/2)} \right]^{2} + \left[ \frac{s_{1} - s_{2}}{2\sin(\psi/2)} \right]^{2} + 1 \right\}^{-1/2}$$
(4.9)

for acute angle. Since Equation (4.7) represents a plane, the normal line defined by (4.8) can be a family of lines. Let the particular normal line be chosen to be such that it intersects z-axis. The line which is perpendicular to this normal line and yet passes through the origin lies in the plane defined by Equation (4.7) and, from Equation (4.9)

$$\sin \gamma_{t} = \cos \gamma_{n} = \left\{ \left[ \frac{s_{1} + s_{2}}{2\cos(\psi/2)} \right]^{2} + \left[ \frac{s_{1} - s_{2}}{2\sin(\psi/2)} \right]^{2} + 1 \right\}^{-1/2}$$

where  $\gamma_{t}$  is the angle between this line and the z-axis. The slope

(tanS<sub>p</sub>) of this line to the xy plane can, thus, be calculated as

$$\tan S_{R} = \tan \gamma_{n} = \left\{ \left[ \frac{s_{1} + s_{2}}{2\cos(\psi/2)} \right]^{2} + \left[ \frac{s_{1} - s_{2}}{2\sin(\psi/2)} \right]^{2} \right\}^{1/2}$$

$$= (s_{1}^{2} + s_{2}^{2} - 2s_{1} s_{2} \cos\psi)^{1/2} / \sin\psi \qquad (4.10)$$

for acute angle. The equation for this normal line, from Expression (4.8), is

$$\begin{bmatrix}
\frac{x - x_0}{s_1 + s_2} \\
\frac{1}{2\cos(\psi/2)}
\end{bmatrix} = \frac{y - y_0}{-\left[\frac{s_1 - s_2}{2\sin(\psi/2)}\right]} = z - z_0$$

For  $(x_0, y_0, z_0) = (0, 0, z_0)$  which is a point on the z-axis,

$$\frac{-x}{\left[\frac{s_1 + s_2}{2\cos(\psi/2)}\right]} = \frac{y}{-\left[\frac{s_1 - s_2}{2\sin(\psi/2)}\right]} = z - z_0$$
 (4.11)

Equation (4.11) can be written also as

$$\left[\frac{s_1 - s_2}{2\sin(\psi/2)}\right] x + \left[\frac{s_1 + s_2}{2\cos(\psi/2)}\right] y = 0$$
 (4.12)

$$x - \left[\frac{s_1 + s_2}{2\cos(\psi/2)}\right] (z - z_0) = 0$$
 (4.13)

$$y + \left[\frac{s_1 - s_2}{2\sin(\psi/2)}\right] (z - z_0) = 0$$
 (4.14)

Equation (4.12) is, indeed, the direction of the line on the xy-plane which is perpendicular to Equation (4.11) and has a slope shown by Equation (4.10) with the xy-plane. Therefore, from Equation (4.12),

$$y = \left[\frac{s_2 - s_1}{s_2 + s_1} \cot(\psi/2)\right] x$$

or the angle  $\theta$  with the x-axis is

$$\tan\theta = \frac{s_2 - s_1}{s_2 + s_1} \cot(\psi/2) \tag{4.15}$$

The whole mathematical scheme so far can be categorized as:

- (1)  $\overrightarrow{OA}_1$  represents the slope vector for the ascending orbit which intersects decending orbit at an intersecting ground angle of  $\psi$ . The decending orbit has the slope vector  $\overrightarrow{OA}_2$ . The reference system is the local Rectangular Cartesian Coordinates with x-axis in East, y-axis in North and z-axis in the direction of the local gravity. There is no separation between  $\overrightarrow{OA}_1$  and  $\overrightarrow{OA}_2$  at the origin 0.
- (2) Equation (4.7) represents the plane which contains the slope vectors  $\overrightarrow{OA}_1$  and  $\overrightarrow{OA}_2$ .
- (3) The true local slope vector of the plane represented by Equation (4.7) has the magnitude tanS, represented by Equation (4.10), and has the true directional angle  $\theta$  with the x-axis as represented by Equation (4.15). This angle  $\theta$  has  $180^{\circ}$  ambiguity. Nevertheless it should not be difficult to figure out the correct one.

It is essential, and, also, important to evaluate the effects of the noises in the measurements of  $s_1$  and  $s_2$  on Equations (4.10) and (4.15). Let

$$s_1 = s_1 + 2\Delta s_1$$
 (4.16)

$$\mathbf{s}_2 = \mathbf{S}_2 + 2\Delta \mathbf{S}_2 \tag{4.17}$$

where  $S_1$  and  $S_2$  are the desired measurements and  $\Delta S_1$  and  $\Delta S_2$  are the noises associated with the measurements of  $S_1$  and  $S_2$ , respectively. The factor 2 is for the worst case which indicates the same noise at the origin with opposite sign. In space application it will be further assumed that the noise levels on the measurements on each track are the same, i.e.,

$$\Delta S_1 = \Delta S_2 = \Delta S \tag{4.18}$$

This is not generally true, but serves to bound the problem.

And in Equation (4.18),  $\Delta S$  is the Root-Mean-Square (RMS) noise in the measurement. Therefore, Equations (4.16) and (4.17) are

$$S_1 = S_1 + 2\Delta S$$
 (4.19)

$$S_2 = S_2 + 2\Delta S$$
 (4.20)

Henceforth, from Equations (4.10), (4.19) and (4.20)

$$tanS = [(S_1 + 2\Delta S)^2 + (S_2 + 2\Delta S)^2 - 2(S_1 + 2\Delta S)$$

$$\cdot (S_2 + 2\Delta S) \cos \psi]^{1/2} / \sin \psi$$
(4.21)

and, from Equation (4.15), (4.19) and (4.20)

$$\tan\theta = \frac{s_2 - s_1}{s_2 + s_1 + 4\Delta s} \cot \psi \tag{4.22}$$

In general,  $S_1$  and  $S_2$  do not equal to each other. And, furthermore,  $S_1 \not> 2\Delta S$  and  $S_2 \not> 2\Delta S$ , Equations (4.21) and (4.22) appear to be highly nonlinear and should be treated accordingly for the effects of  $\Delta S$  on S and  $\theta$ . Nevertheless, after regrouping

$$\tan S = \left[\tan^2 S_R + \tan^2 \Delta S_{\varepsilon} + \tan^2 \Delta S_{\varepsilon}' - (S_1 + S_2 - 2\Delta S)^2 / \sin^2 \psi\right]^{1/2}$$

$$(4.23)$$

and

$$tan\theta = [1 - F(S_1, S_2, \Delta S)] tan\theta_R$$
 (4.24)

where

$$\tan S_{R} = \left[ \frac{S_{1}^{2} + S_{2}^{2} - 2S_{1}S_{2}\cos\psi}{\sin\psi} \right]^{1/2}$$
 (4.25)

$$\tan\Delta S_{\varepsilon} = \frac{\left[2(2\Delta S)^{2} - 2(2\Delta S)^{2}\cos\psi\right]^{1/2}}{\sin\psi}, \qquad (4.26)$$

$$\tan \Delta S_{\varepsilon} = \frac{\left[ (2\Delta S)^{2} + (S_{1} + S_{2})^{2} - 2(2\Delta S)(S_{1} + S_{2})\cos \psi \right]^{1/2}}{\sin \psi},$$
(4.27)

$$\tan \theta_{R} = \frac{S_{2} - S_{1}}{S_{2} + S_{1}} \cot \psi , \qquad (4.28)$$

and

$$F(S_1, S_2, \Delta S) = \frac{2(2\Delta S)}{S_2 + S_1 + 2(2\Delta S)}$$
 (4.29)

The terms  $\tan S_R$  and  $\tan \theta_R$  are the desired terms without the effects of the measurement noise. The terms  $\tan \Delta S_{\epsilon}$ ,  $\tan \Delta S_{\epsilon}$ , and  $F(S_1, S_2, \Delta S)$  are all introduced by non-zero  $\Delta S$ .

In view of the nonlinear characters in terms of  $S_1$ ,  $S_2$ , and  $\Delta S$  in Equations (4.23) and (4.24) for S and  $\theta$  it is desirable to set  $S_1$  and  $S_2$  to zero in order to demonstrate the different effects on S and  $\theta$  due to the same  $\Delta S$  with various orbit inclinations. Thus, Equations (4.25), (4.26), (4.27), (4.28), and (4.29) are reduced to

$$tanS_p = 0 , (4.30)$$

$$tan\Delta S_{\varepsilon} = \frac{2\sqrt{2} \Delta S (1 - \cos \psi)^{1/2}}{\sin \psi}, \tag{4.31}$$

$$tan\Delta S_{\varepsilon}' = \frac{2\Delta S}{\sin \psi} , \qquad (4.32)$$

$$tan\theta_{p} = 0 , (4.33)$$

and

$$F(0, 0, \Delta S) = 1$$
 (4.34)

Therefore Equations (4.23) and (4.24) become

$$tanS = tan\Delta S_{\epsilon} = \frac{2\sqrt{2} \Delta S (1 - cos\psi)^{1/2}}{sin\psi}$$

=  $2\Delta S \sec(\psi/2)$  , (4.35)

and

 $tan\theta = tan\theta_{p} = 0 (4.36)$ 

Equations (4.35) and (4.36) show that the local slope generated by  $\Delta S$  alone is in either the east or the west direction. The conclusions presented by Equations (4.35) and (4.36) are quite obvious and should be used to ascertain the validities of Equations (4.23) and (4.24).

Nevertheless, in a typical Western North Atlantic Ocean, let  $S_1$  = -1 sec and  $S_2$  = -9/10 sec which give a Gulf-Stream-like current. Let the root-mean-square noise be 1 sec, i.e., 2.5 cm root-mean-square altimeter precision with measurements separated by 10 km [ref. Figure 5.1, Equations (4.16) and (4.17)]. By using Equations (3.1), (4.23), (4.24), (4.25), (4.26), (4.27), (4.28) and (4.29), for the worst possible case where  $\Delta S$  assumes the same sign as those of  $S_1$  and  $S_2$  Table 4.1 can be computed for orbit inclinations 93°, 98° and 108° where

- $\psi$  = crossing angle between ascending and descending orbits as defined in Figure 4.1 and given in Figure 4.2,
- $V_p$  = the real geostrophic current speed in m/sec,
- V = The geostrophic current speed in m/sec computed with the non-zero root-mean-square noise,
- $\delta_{
  m R}$  = the real direction of geostrophic current with respect to the local east in a counter clockwise direction, and
- $\delta$  = the direction derived for the geostrophic current V in the same sense of direction as defined for  $\delta_{\rm p}$  .

Figure 4.3 shows the plots of  $V_R$  and V with respect to the latitude for orbit inclinations of 93°, 98°, and 108° while Figure 4.4 shows the plots of  $\delta_R$  and  $\delta$  for these  $V_R$  and V. The broken lines in Figures 4.3 and 4.4 represent V and  $\delta$ , respectively, which are computed with the Root-Mean-Square (RMS) noise  $\Delta S$ , and the solid lines represent  $V_R$  and  $\delta_R$ , respectively, which are, of course, computed without the RMS noise.

 $V_R$ , V,  $\delta_R$ , and  $\delta$  computed with  $S_1 = -1$  sec.,  $S_2 = -0.9$  sec., and  $\Delta S = 1$  sec for 93°, 98°, and 108° Orbit Inclinations

	80	144.92 3.079 9.634 4.29 1.37 1.022 3.197 94.29	73.46 1.211 3.626 -0.92 -0.29 0.40 1.20 89.08	
	20	162.4 6.064 19.011 9.43 3.03 2.109 7.174 99.43	131.98 2.290 7.144 2.69 0.86 0.80 2.48 92.69 90.86	50.76 1.143 3.219 -2.46 -0.79 0.40 1.12 87.54 89.21
	09	167.98 8.858 27.779 13.87 4.57 3.343 10.483 103.87 94.57	147.68 3.333 10.431 4.74 1.55 1.26 3.94 94.74	103.66 1.523 4.701 0.74 0.23 0.57 1.77 90.74
	50	170.66 11.397 35.748 17.74 5.88 4.862 15.250 107.74 95.88	155 4.28 13.409 6.45 2.06 1.83 5.72 96.45	122.54 1.943 6.045 1.95 0.63 0.83 2.58 91.95
	40	172.16 13.589 42.624 20.91 7.01 6.909 21.670 110.91	159.06 5.106 16.004 7.86 2.52 2.60 8.14 97.86	132.42 2.309 7.205 2.75 0.86 1.17 3.66 92.75
	30	173.08 15.408 48.332 23.46 7.97 10.071 31.590 113.46	161.5 5.770 18.088 8.92 2.92 3.77 11.82 98.92	138.18 2.604 8.137 3.38 1.09 1.70 5.32 93.38
	20	173.62 16.661 52.264 25.22 8.64 15.920 49.938 115.22 98.64	162.96 6.255 19.610 9.76 3.15 5.98 18.73 99.76	141.6 2.826 8.837 3.78 1.20 2.70 8.44 93.78
	10	173.9 17.447 54.729 26.20 8.98 32.835 102.998 116.2 98.98	163.76 6.552 20.542 10.26 3.32 12.33 38.65 100.26 93.32	143.42 2.595 9.256 4.06 1.32 5.57 17.42 94.06 91.32
	0	174 17.787 55.796 26.61 9.15	164 6.648 20.844 10.43 3.38 - -	144 3.004 9.398 4.12 1.32
Latitude	Orbit Inclination	tan S tan S θ R V V δ R	ψ tan S tan S θ θ V V V δ δ δ	ος κα ος κα ο
	Orbit Incli	93.	.86	108°



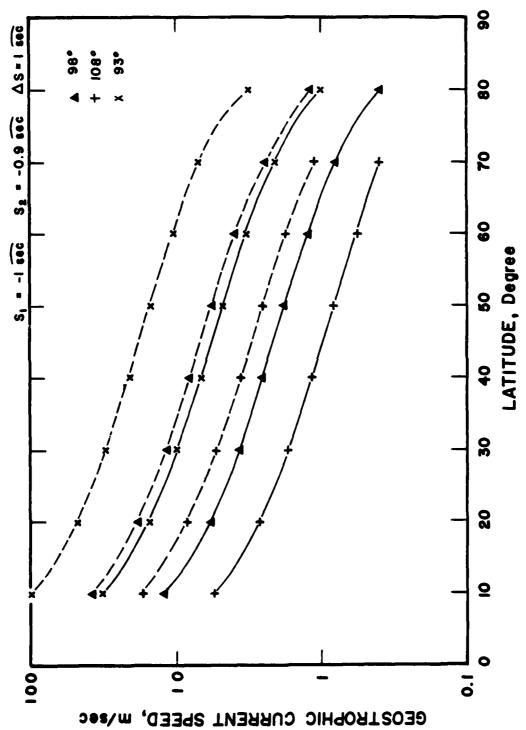


Fig. 4.3 Geostrophic currents inferred w/ (dashed curves) and w/o (solid curves) RMS noise for  $93^{\rm O}$ , and  $108^{\rm O}$  orbit inclinations

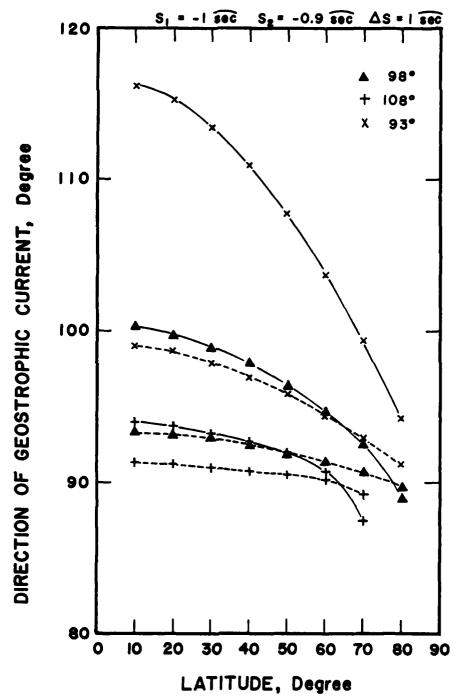


Fig. 4.4 - Geostrophic current direction wrt the east w/ (dashed curves) and w/o (solid curves) RMS noise for  $93^{\circ}$ ,  $98^{\circ}$ , and  $180^{\circ}$  orbit inclinations

However, several observations can be made regarding Figures 4.3 and 4.4. (1) There are nearly constant relationships resulting from the assumptions such as

$$V_{R} = C V \tag{4.37}$$

that can be made between each pair of  $V_R$  and V for various orbit inclinations. In reality, the value of  $\Delta S$  is Gaussian-distributed, and a range of C in Equation (4.37) should be evaluated, for each orbit inclination, and independently verified by selective ground truth programs. Nevertheless orbit inclination  $108^\circ$  does have the smallest magnitude for C among the inclinations selected. (2) The effects due to RMS noise on  $\delta_R$  appear to bend  $\delta_R$  toward local north or south in direction as predicted by Equation (4.36). (3) The orbit inclination  $108^\circ$  shows to be the best among the three orbit inclinations in the determination of reasonable  $V_R$  and  $\delta_R$ . Therefore, this orbit inclination is the best for geostrophic current determination for the choices selected.

### 5.0 UNCERTAINTY IN SATELLITE ALTIMETER MEASUREMENT OF OCEAN SURFACE SLOPE

A slope of  $10^{-5}$  (10 cm height difference in 10 km) is approximately equal to an angular slope of  $2\sqrt{\text{sec}}$ . If all of the assumptions of Section 2.1 are satisfied, the uncertainty of the altimeter along-track slope determination is limited by the measurement precision.

Conservative expressions for along-track slope measurement uncertainties are shown in Figure 5.1 for the cases of 5 cm RMS and 10 cm RMS altimeter noise  $\Delta S$ . The RMS slope uncertainty is expressed in dimensions of sec. In the estimation, if the RMS noise is 5 cm, and the measurements are made at 2 points separated by 10 km, the RMS slope uncertainty is  $10^{-5}$ , or 2 sec, which is for the worst case estimation.

It has been demonstrated that proper filtering (weighted averaging) of the along-track radar altimeter data can greatly reduce the along-track slope measurement uncertainties. Uncertainties as low as 0.1 sec have been obtained from SEASAT altimeter data (B. Chovitz and B. Douglas, National Oceanic Survey, NOAA, private communications). The averaging techniques define the space-scale of precise slope determinations.

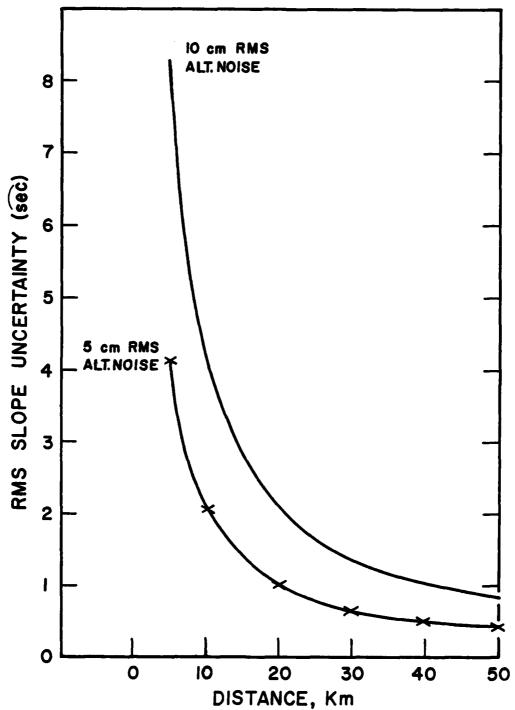


Fig. 5.1 - Along-track slope measurement uncertainties

Since the current-speed uncertainties are directly proportional to slope uncertainties, without elaborated discussions on the proper weighting techniques and the spatial scale factors, illustrative examples of current speed determinations are based upon the conservative expressions as shown in Figure 5.1 wherever they are appropriate.

The uncertainty of determination of the components of surface slope along-track may be expressed as a product of a latitude-inclination-direction-dependent Error Factor with the along-track slope uncertainty. The latitude-dependence of North and East slope error factors for selected orbit inclinations of 73°, 100°, 105°, and 110° are shown in Figure 5.2.

The slope component uncertainties as a function of latitude for an along-track slope uncertainty of 0.85 sec and an orbit inclination of 98° (approximately sun-synchronous) are shown in Figure 5.3. Under the conservative assumptions of Figure 5.1, a slope uncertainty of 0.85 sec implies a distance separation of 50 km for 10 cm RMS altimeter noise, or a distance separation of 25 km for 5 cm RMS altimeter noise.

### 6.0 DISCUSSION

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From Figures 4.3 and 4.4 it is seen that: (1) The uncertainties in the determination of geostrophic current rapidly decrease in their magnitude and, yet, maintain the nearly constant relationship as given by Equation (4.37) for increasing values of latitude  $\phi$ . Nevertheless, a range of values of C, in Equation (4.37), can be determined by ground truth for Gaussian-distributed RMS noise  $\Delta S$ . For each orbit inclination the distribution of C and its lower and upper bounds for the values of C can be calculated. With the known C,  $\Delta S$ , V, and  $\delta$ , the true local geostrophic current parameters  $V_R$  and  $\delta_R$  can be evaluated by Equation (4.37) with Figures 4.3 and 4.4. (2) The effects due to RMS noise on the determination of geostrophic current directions show the directional angle  $\delta$ , with respect to the local east, to be more in local east—west oriented direction than that shown by the true directional

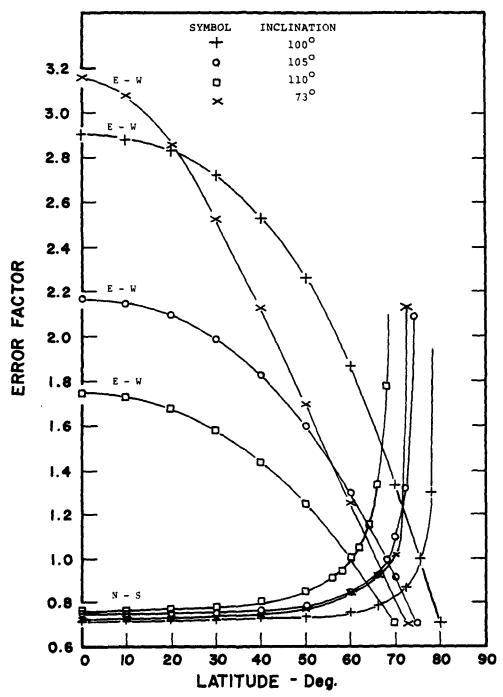


Fig. 5.2 - Error factors in N - S and E - W directions for different orbit inclinations

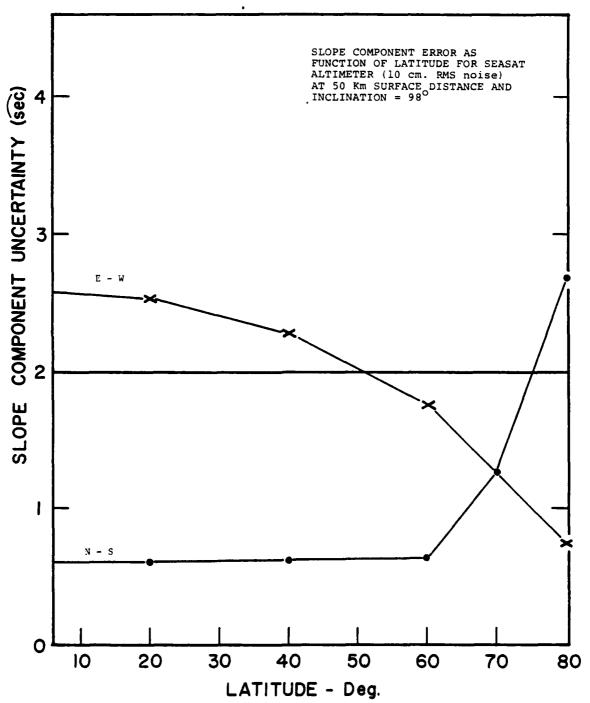


Fig. 5.3 - Along-track slope component uncertainties with 0.85  $\sec$  slope uncertainty for 98° orbit inclination

angle  $\delta_R$ . (3) For  $S_1 = -1$  sec,  $S_2 = -0.9$  sec, and  $\Delta S = 1$  sec the best orbit inclination, among the three studied for the detection of geostrophic current velocity, is  $108^\circ$ . The second choice appears to be  $98^\circ$ . (4) It is also important to point out that, with the assumptions made for the derivation for geostrophic current calculation, the results are meaningful only for latitude  $20^\circ$  and up.

As has been demonstrated from SEASAT-A analyses, an altimeter precision of 10 cm is adequate to detect and provide first-order speed estimates for Western Boundary currents (e.g. Gulf Stream, Kuroshio) which have characteristic speeds of the order 1 m/sec, and represent the higher portion of the range of current speeds to be measured. The precision goal for measurement of current speeds within the National Oceanic Satellite System (NOSS) is 5 cm/sec. From Figures 4.3 and 4.4 and Equations (3.3), (3.3a), and (3.3b), it is seen that altimeter surface measurement uncertainty must be of the order of 0.1 sec to achieve the NOSS current speed measurement goal for 20° latitude and up.

The altimeter measurement requirement for 0.1  $\sec$  precision is equivalent to a slope precision of  $5\times10^{-7}$  (5 cm in 100 km). This slope precision is necessary to determine the speed of such important features as the Southern Ocean Current and the Laborador Drift Current.

Since in reality, the assumptions made in Section 2.1 cannot be fully satisfied, the averaging and optimizing techniques such as those employed by Huang, et al (1978) should be used so that the effects due to the partial invalidaties of some of the assumptions can be alleviated.

For high latitudes a network of in-situ current measurements used by Wunsch (1977) can certainly be used to reduce the uncertainties in the altimeter-derived current measurements. These uncertainties are primarily, caused by the small crossing angles and the lack of information on the north component of the slopes at the high latitudes. The main purpose of the in-situ current measurement is to provide the boundary condition for the current system.

A multi-beam altimeter (McGoogan and Walsh, 1978) can provide several off-nadir measurements simultanteously. The deployment of this type of instrument along-track as well as across-track in space will have several advantages:

- (1) The true local ocean surface slope vectors can be obtained for the individual track without waiting for crossings to occur. Henceforth, the true local geostrophic ocean current velocities, not only the components across-track, can be inferred from a single pass.
- (2) The effects due to orbit altitude rate uncertainties on the ocean surface slope measurements are greatly reduced in comparison to those provided by the single-beam nadir altimeter.
- (3) The resolution of both across-track and along-track measurements can be changed if desired.

Further airborne tests and analyses of this multi-beam instrument are required to be pursued in order to answer some of the questions associated, always, with new instruments.

### 7.0 CONCLUSION

Orbit inclination of 108° is the best choice among the three selected orbit inclincations (93°, 98° and 108°) for the detection of geostrophic current vectors above 20° latitude. The second choice is 98° orbit inclination. The relationship shown by Equation (4.37) as

$$v_R = c v$$

where  $\mathbf{V}_{\mathbf{R}}$  is the true local geostrophic current speed, V is the geostrophic current speed evaluated with the RMS noise, and C is a coefficient can be used to study the possible accuracy of recovering geostrophic current velocities from satellite radar altimeter measurements

of varying precision and in different orbits.

The uncertainty in geostrophic current velocity determination is a function of the altimeter along-track slope measurement precision. Altimeter slope measurement uncertainty must be less than 0.1 sec to achieve a minimum uncertainty in geostrophic current speed determination of 5 cm/sec.

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